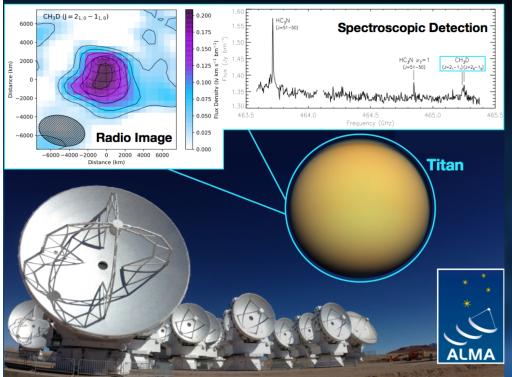
## Measuring Titan's Methane from Earth



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(Bottom) The ALMA telescope in Chile. (Middle) Saturn's largest moon, Titan, as seen from the Cassini spacecraft. (Top left): A radio image of Titan's  $CH_3D$  taken with ALMA. (Top right): The submillimeter spectrum of  $CH_3D$ . (Images from NRAO/ESA/NAOJ; NASA/JPL-Caltech/SSI; this work.) Following the end of the successful Cassini/Huygens mission, ground-based observations in the submillimeter wavelength range are now helping to answer outstanding questions on the composition and variability of Titan's complex atmosphere.

- Saturn's largest moon, Titan, has a substantial atmosphere mostly composed of methane (CH<sub>4</sub>) and molecular nitrogen (N<sub>2</sub>). These gases react with ions and solar radiation to create a wealth of organic trace constituents.
- Many of Titan's atmospheric species vary with latitude and time throughout Titan's ~29.5 year seasonal cycle. However, variations in methane are still poorly understood, despite its importance in Titan's atmospheric chemistry.
- We have detected an isotope of CH<sub>4</sub> (CH<sub>3</sub>D) in Titan's atmosphere using the Atacama Large Millimeter/submillimeter Array (ALMA), a telescope comprised of 66 individual antennas in the Atacama Desert in Chile. This is the first detection of CH<sub>3</sub>D at submillimeter wavelengths, and allows for monitoring of Titan's global methane distribution, and its seasonal variability in the post-Cassini era.